



## Domain Feature Enhanced Unisolver for Weak Lensing Inference

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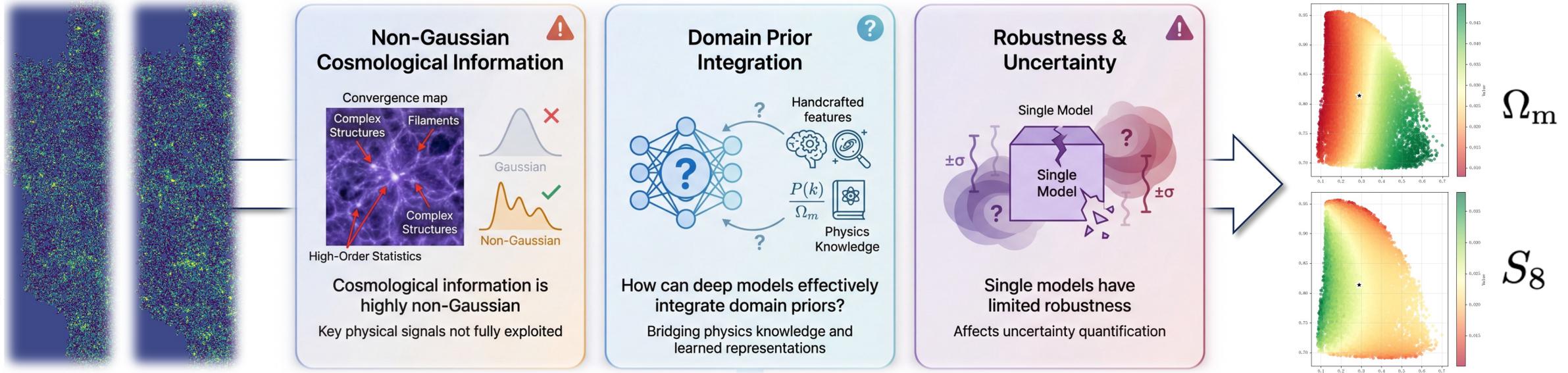


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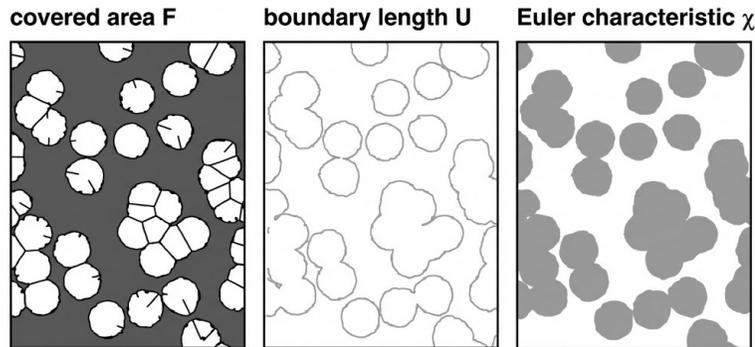
# Challenges in Weak Lensing Inference



- 1. Scientific problem:** requires capturing physical information behind observation
- 2. Learning difficulty:** effectively integrate both deep representations and domain priors
- 3. Ill-posed problem:** inferring underlying coefficients from partial observations

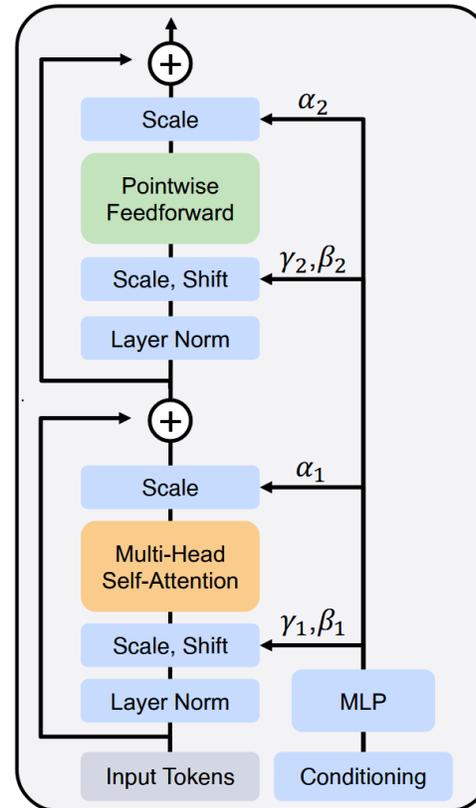
# Key Designs in Our Method

## (1) Elaborated Domain Priors

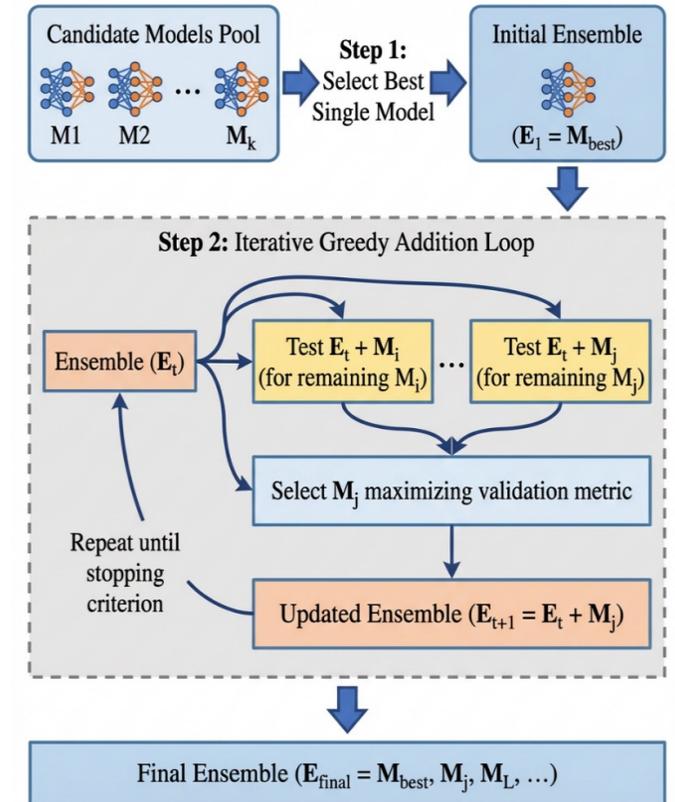


- Power Spectrum
- Peak Counts
- Minkowski Functionals
- Multiscale Wavelet Statistics
- 1-point PDF
- ... ..

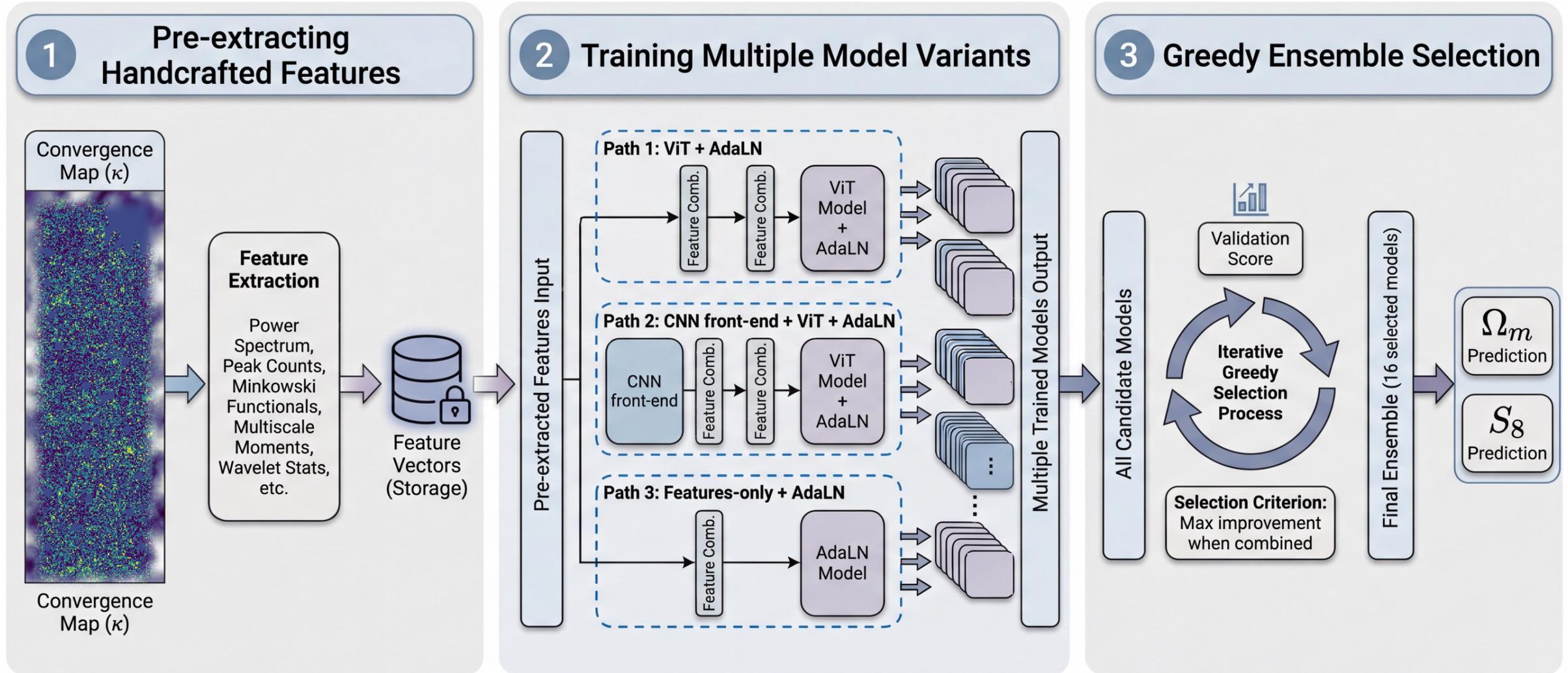
## (2) Specialized Conditional Transformer



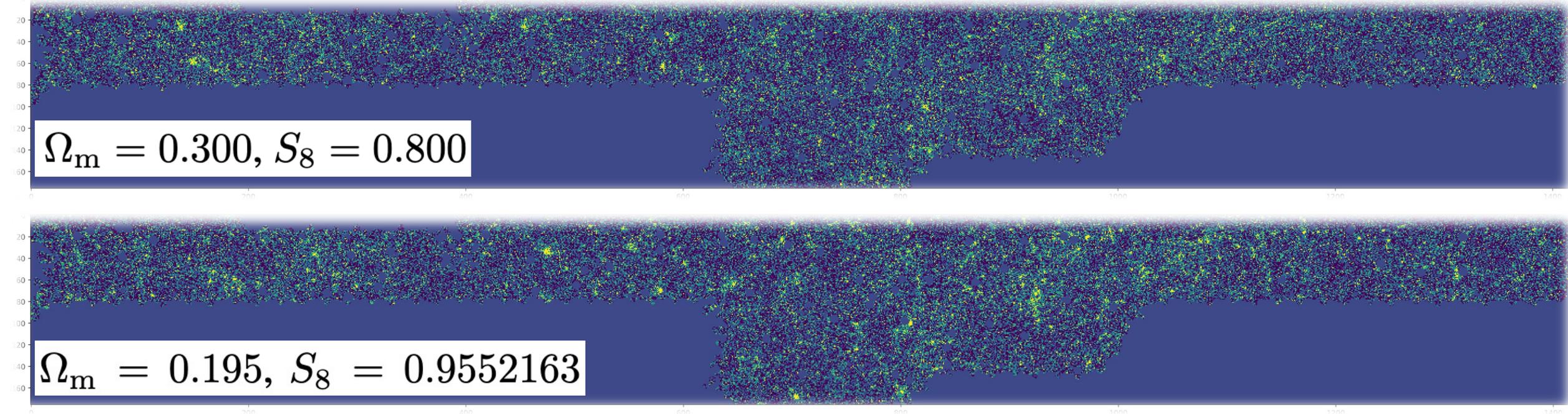
## (3) Robust Ensemble Algorithm



# Overall Training and Inference Pipeline



# Stage 1: Starting From Domain Priors

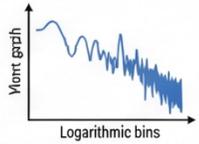


- Weak lensing is a **highly structured physics signal**, not arbitrary images.
- Domain-specific priors are based on **decades of accumulated knowledge** in the field.

**Effectively capture information from low signal-to-noise ratio cosmological observations**

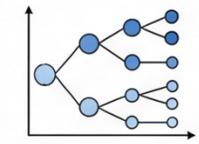
# Stage 1: Domain Features (11 Categories)

## Spectral & Frequency Features



### Power Spectrum

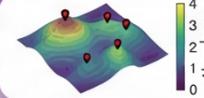
Second-order statistics, Gaussian modes, logarithmic multipole bins



### Multiscale Wavelet Statistics

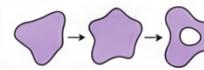
Amplitude, variation, higher-order stats at different scales

## Morphological & Topological Features



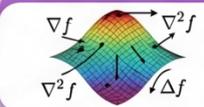
### Peak Counts

Local maxima abundance, non-Gaussian information



### Minkowski Functionals

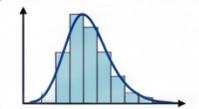
$V_0, V_1, V_2$  - area, perimeter, genus over thresholds



### Gradient and Curvature Statistics

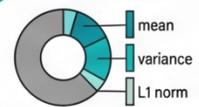
1st & 2nd derivatives, Laplacian, Hessian

## Distribution & Density Features



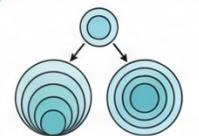
### 1-point PDF

Density fluctuation distribution



### Aperture Mass Statistics

Mean, variance, L1 norm, skewness



### Multiscale Moments

Smoothed moments, edge statistics, multi-resolution

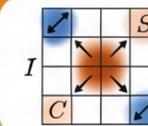
## Statistical & Information Features

### Lightweight Statistical Descriptors

Moments, autocorrelation, spectral slope

### Entropy and Information-Theoretic Features

Fisher info, KL divergence, mutual info



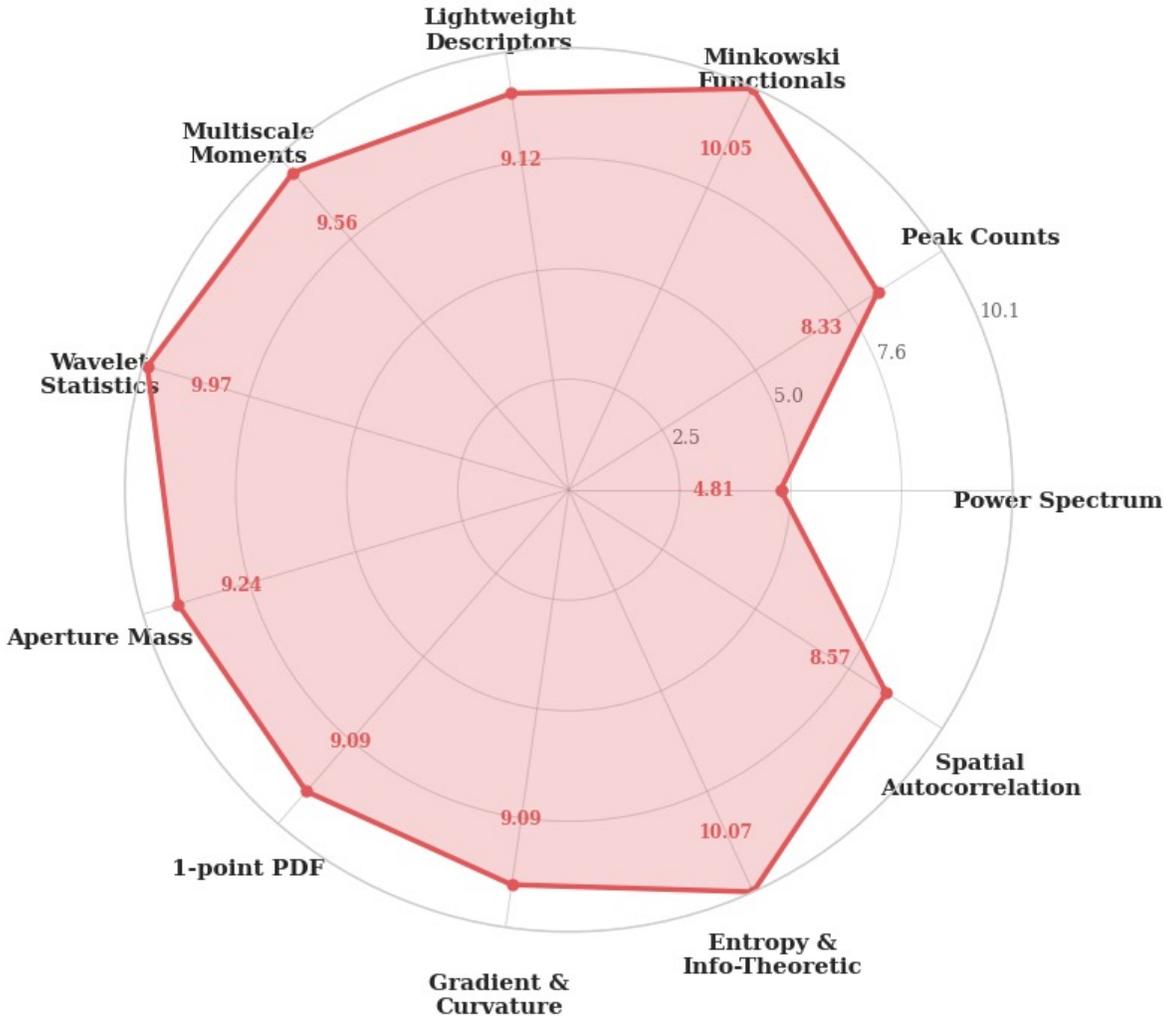
### Spatial Autocorrelation Features

Moran's I, Geary's C, spatial entropy

Domain Features

# Stage 1: Performance of Each Domain Feature

Validation Performance Across Feature Categories

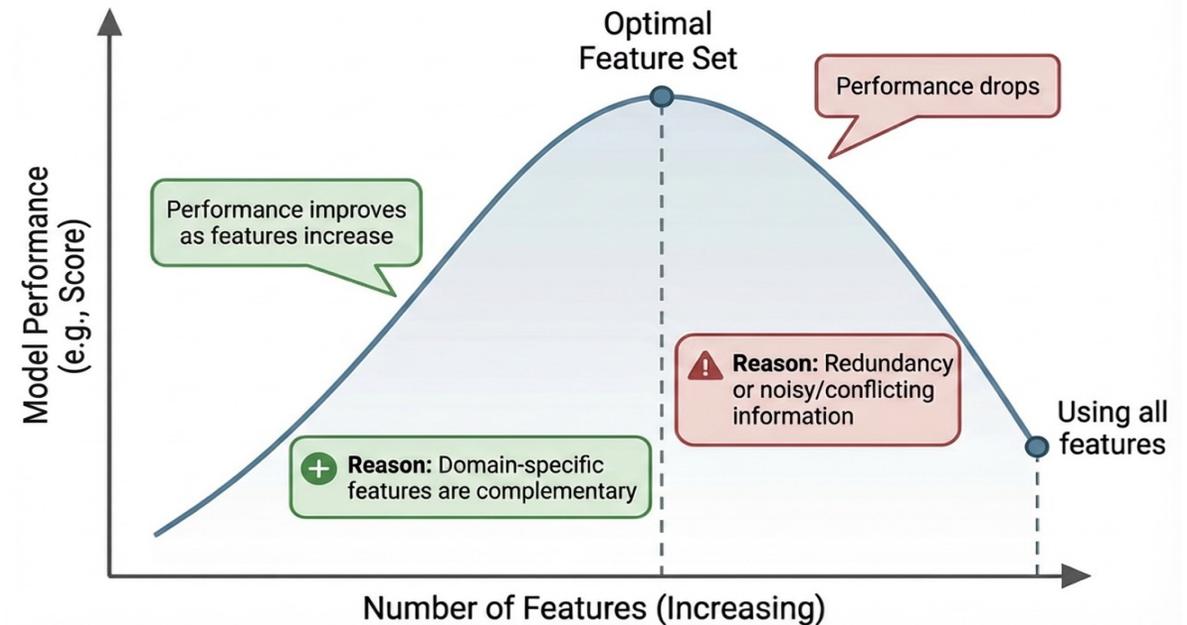


- Splits the public data for train and val as **8:2**.
- Linearly regress the coefficients with MLP.

Feature Name	Val Score
Power Spectrum	4.81
Peak Counts	8.33
 Minkowski Functionals	10.05
Lightweight Statistical Descriptors	9.12
Multiscale Moments	9.56
 Multiscale Wavelet Statistics	9.97
Aperture Mass Statistics	9.24
1-point PDF	9.09
Gradient and Curvature Statistics	9.09
 Entropy and Information-Theoretic Features	10.07
Spatial Autocorrelation Features	8.57

# Stage 1: Complementarity Among Domain Features

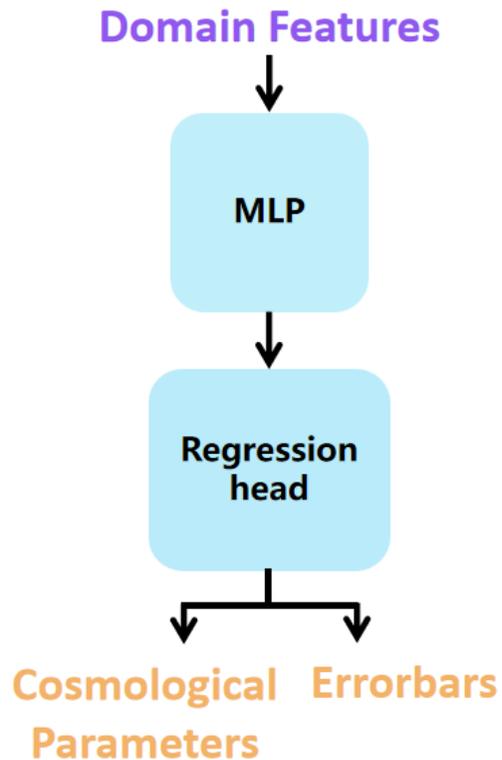
Method	Test Score
CNN	9.30
+Power Spectrum	10.00
+Power Spectrum +Peak Counts	10.31
+Power Spectrum +Peak Counts +Minkowski Functionals	10.95
+Power Spectrum +Peak Counts +Minkowski Functionals +Lightweight Statistical Descriptors	11.08
+Power Spectrum +Peak Counts +Minkowski Functionals +Lightweight Statistical Descriptors +Multiscale Moments	<u>11.24</u>
+All Domain Features	11.22



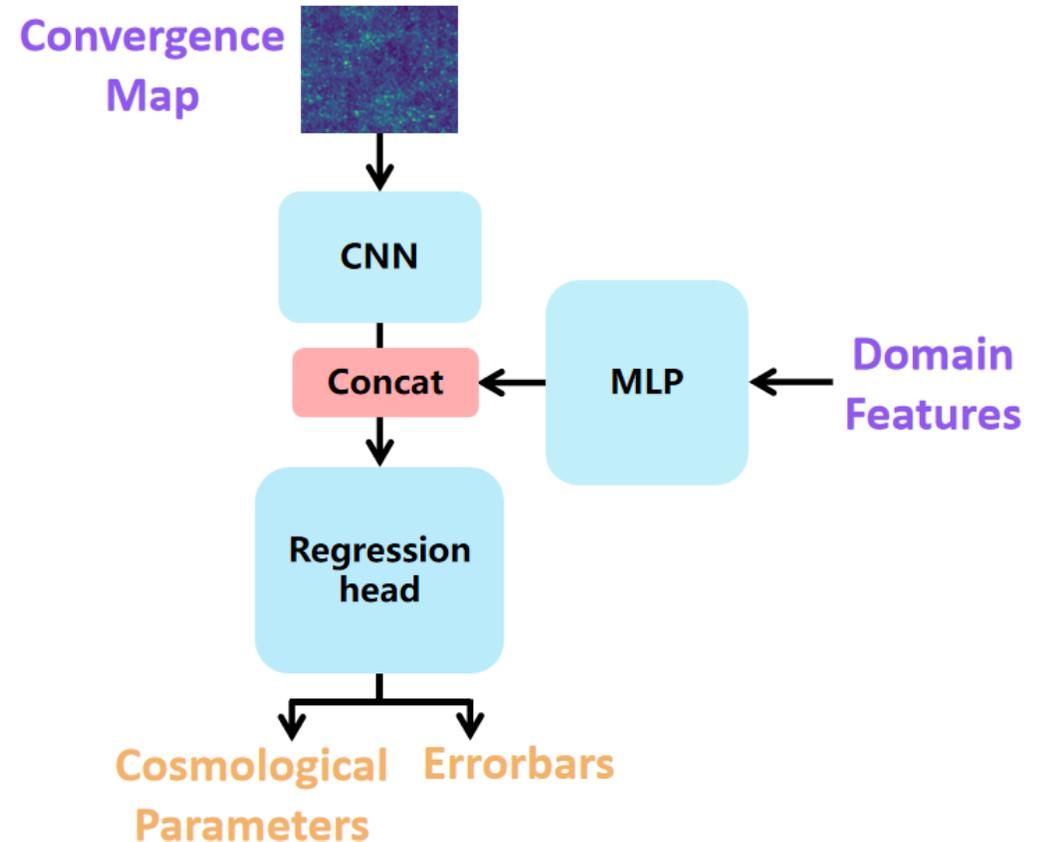
- There exists a **hidden complementarity** among the 11 extracted domain features.
- **But, Too Many Cooks Spoil the Broth.**

# Stage 2: Integrating Deep and Domain features

Architecture 1: Only domain features



Architecture 2: Simple integration in the last layer



The above two simple architectures were explored. **Performance gains were limited.**

# Stage 2: PDE-Conditional Transformer (Unisolver)

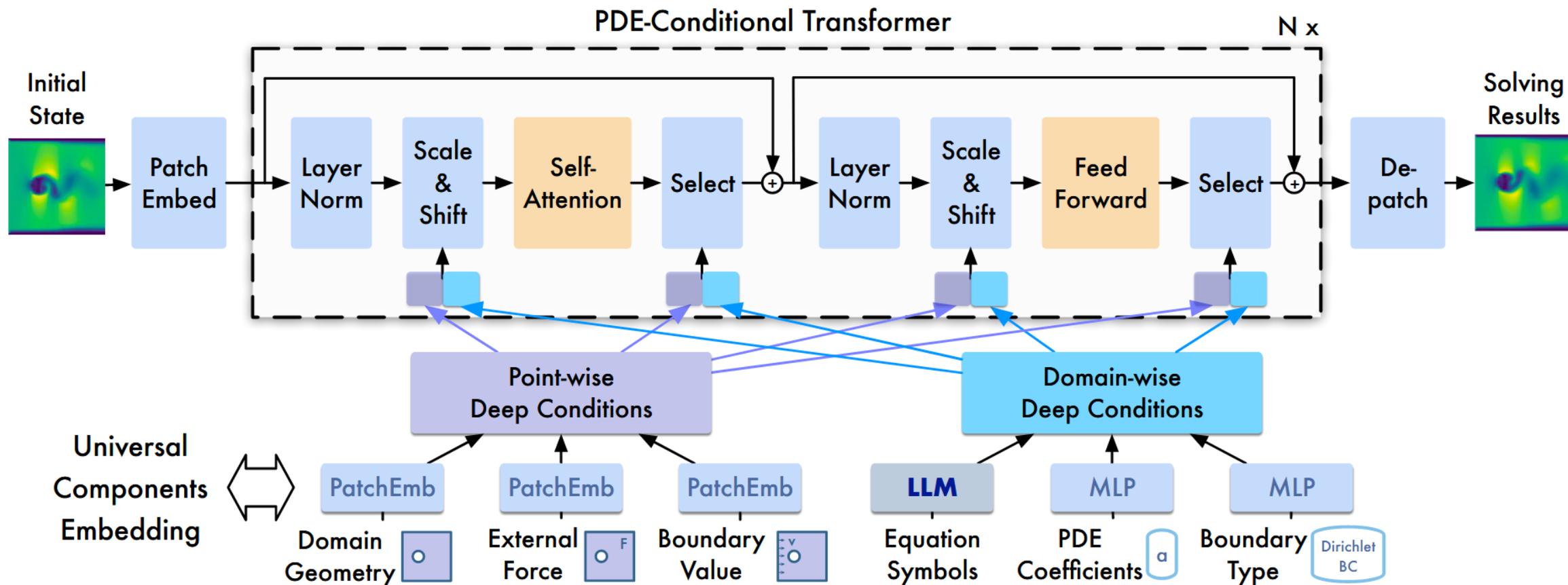


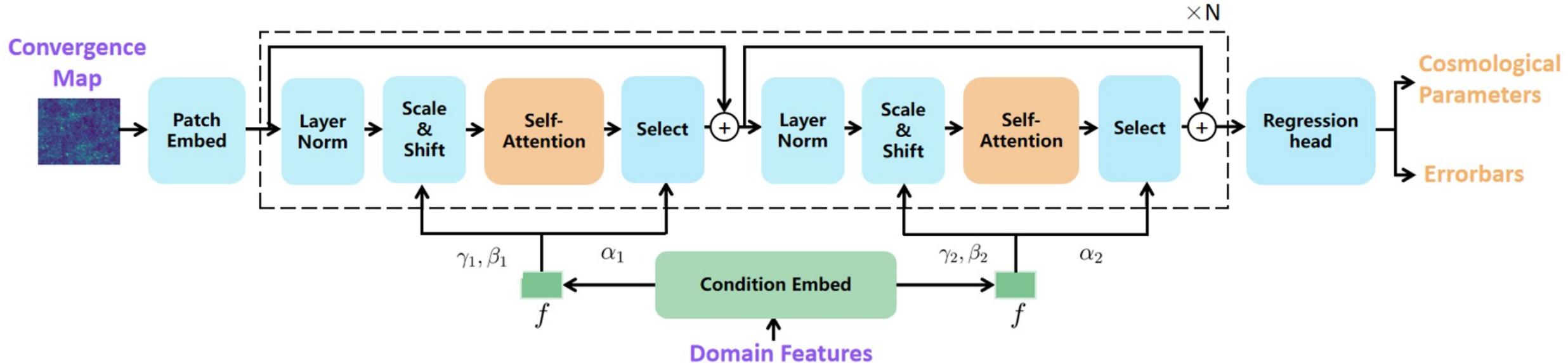
Figure 3. Overview of Unisolver. We universally embed all PDE components, such as domain geometries, equation symbols and boundary types into deep conditions and employ a conditional Transformer to aggregate deep conditions in the decoupled subspace.

# Stage 2: From PDE to Cosmology

Domain-Feature-Conditioned Unisolver

$$\text{AdaLN}(h \mid \underline{f}) = \gamma(f) \cdot \frac{h - \mu(h)}{\sigma(h)} + \beta(f)$$

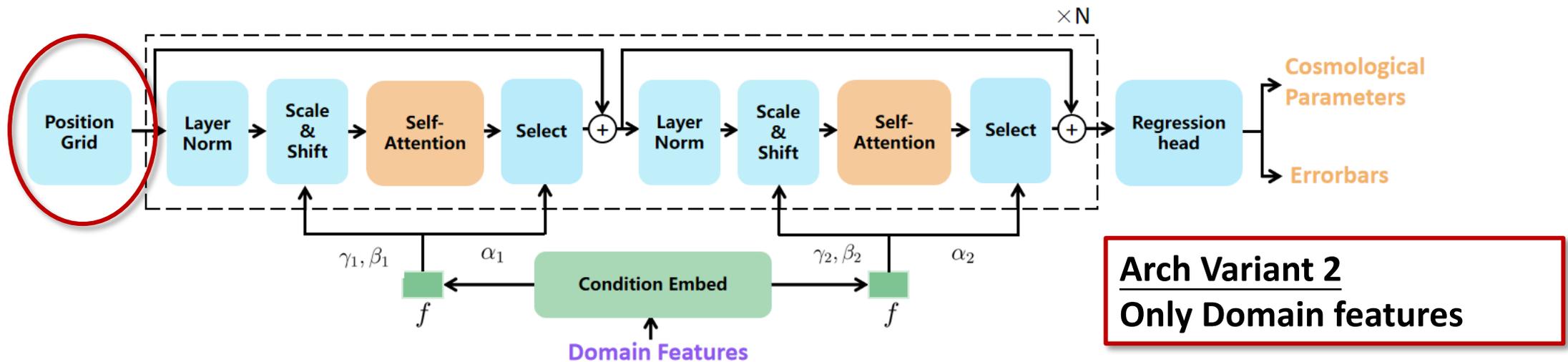
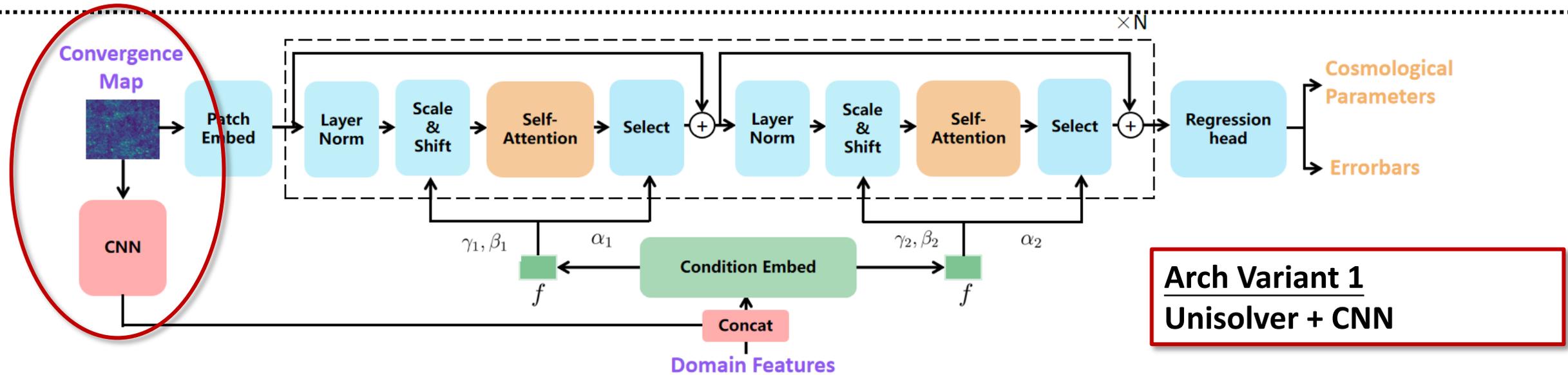
global physical context



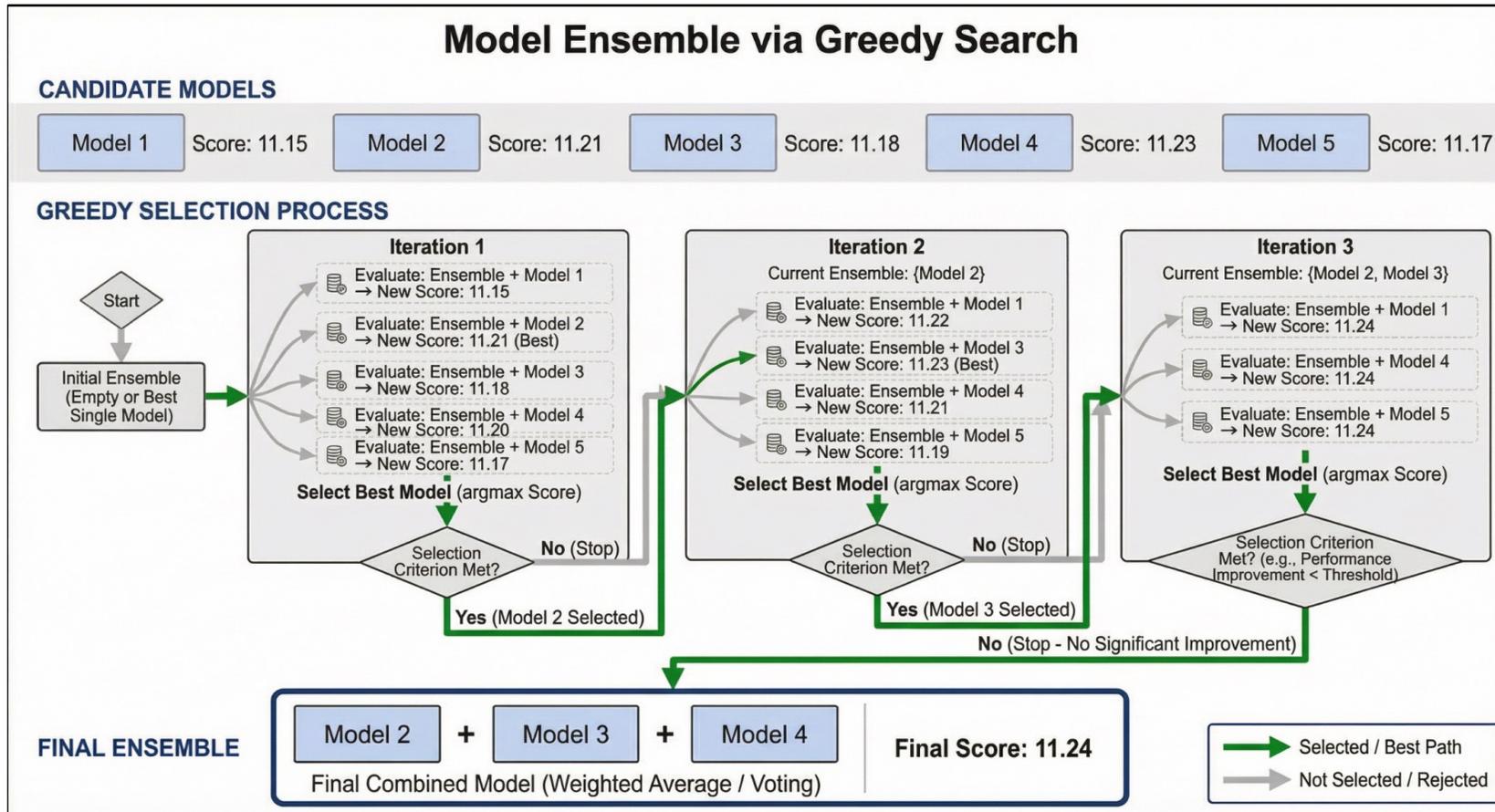
- ✓ From CNN-based to Transformer-based architecture; **Enhanced model capacity!**
- ✓ Adaptively rescale and shift internal representations based on **cosmological conditions**.

# Stage 3: Ensemble Algorithm

**Heterogeneity: (1) Domain Feature Combinations; (2) Architecture Variants.**



# Stage 3: Greedy Model Selection Beyond All-in Ensemble



Huge ensemble Space

Combination of 11 Features

X

3 Architecture Variants

We need to reduce the search complexity

- (Search Order) Adds single models one by one **from high score to low score**;
- (Greedy Ensemble) Keep the newly added model only **if it can improve the performance**.

# Stage 3: Results of the Greedy Search Ensemble

IDs listed under **Handcrafted Feature Combinations** correspond to the following feature groups:

- (1) Power Spectrum,
- (2) Peak Counts,
- (3) Minkowski Functionals,
- (4) Lightweight Statistical Descriptors,
- (5) Multiscale Moments,
- (6) Multiscale Wavelet Statistics,
- (7) Aperture Mass Statistics,
- (8) 1-point PDF,
- (9) Gradient and Curvature Statistics,
- (10) Entropy and Information-Theoretic Features,
- (11) Spatial Autocorrelation Features.

IDs listed under **Model Architecture**

correspond to the following model variants:

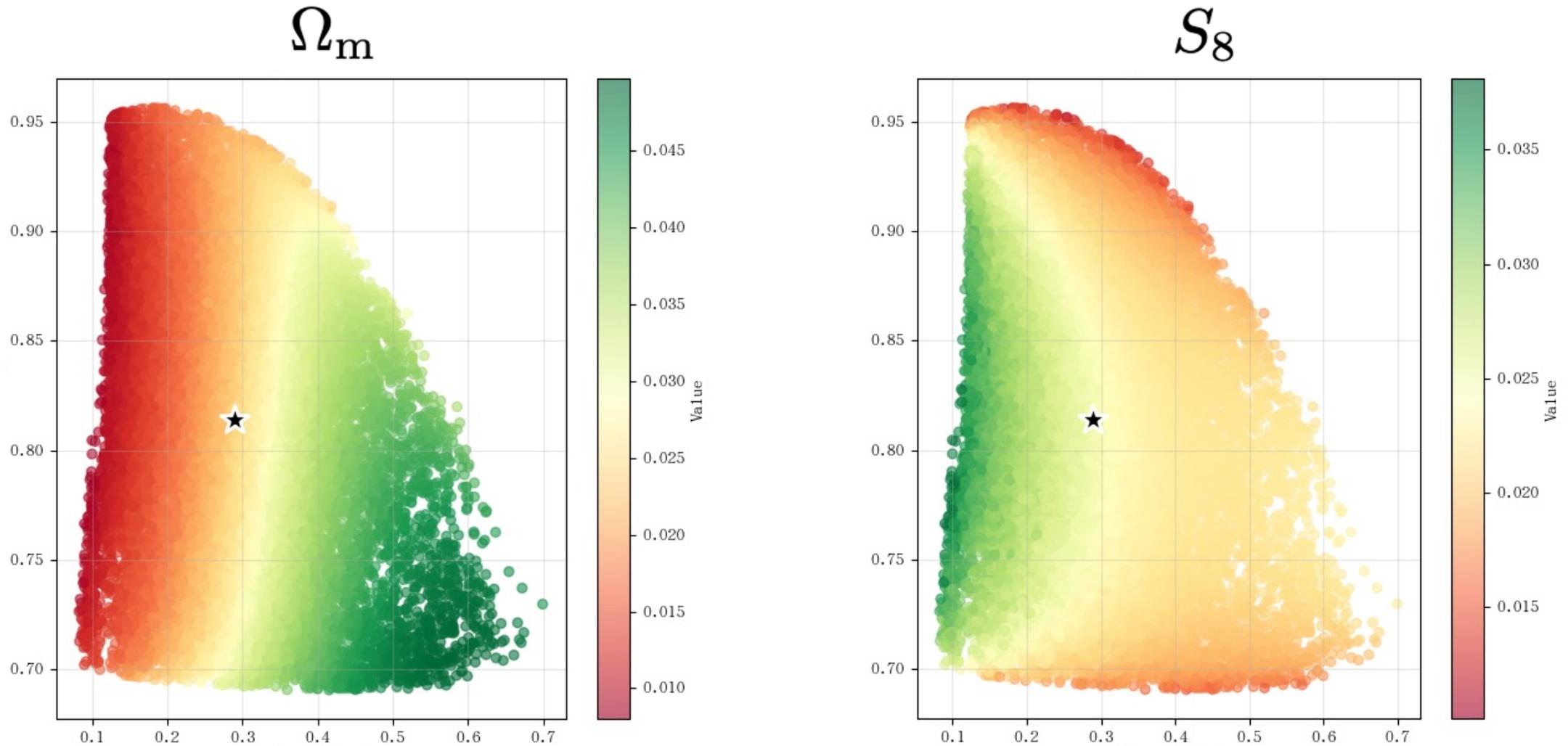
- (1) Standard architecture - ViT with AdaLN,
- (2) Standard architecture with a CNN front-end,
- (3) Standard architecture without convergence map input.

Model ID	Handcrafted Feature Combinations	Model Architecture
1	1,3,4,6,8	1
2	1,4,5,7,8	2
3	1,2,3,6,7	2
4	2,4,5,7,8	2
5	1,2,4,5,7	2
6	1,2,3,4,7	2
7	1,2,3,6,8	2
8	2,3,4,5,7	3
9	2,5,6,7,10	3
10	2,4,5,6,7,10	3
11	2,4,5,7,8	3
12	4,5,7,9,11	3
13	2,4,5,9,10	3
14	2,4,5,9,11	3
15	2,4,8,9,10	3
16	2,5,7,8,9	3

Test Score in the final holdout dataset: **11.2847**

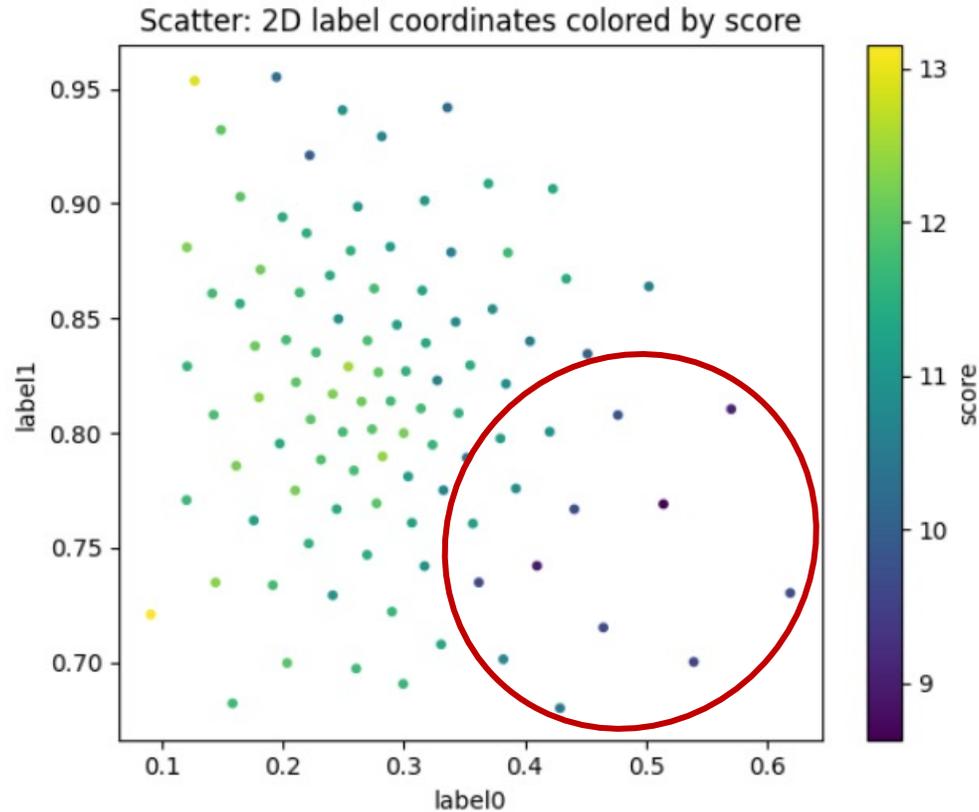
Rank: **3**

# Interesting Finding I — — Errorbar Distribution of Parameters



Errorbar of both parameters exhibits **strong dependence** on  $\Omega_m$ , but **opposite trends**.

# Interesting Finding II — — Score Distribution of Regions



Samples located in the **bottom-right** region of the parameter space tend to exhibit lower scores.

Possible reasons: observation hardness / insufficient data collection.

# Weak Lensing ML Uncertainty Challenge



Thank You!

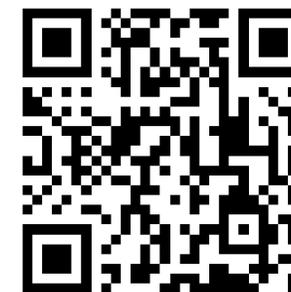
wenghk22@mails.Tsinghua.edu.cn / wuhaixu98@gmail.com

A new technical report is coming soon.



Code for Unisolver

<https://github.com/thuml/Unisolver>



Paper for Unisolver

<https://openreview.net/pdf?id=r1ryQoI9iZ>